



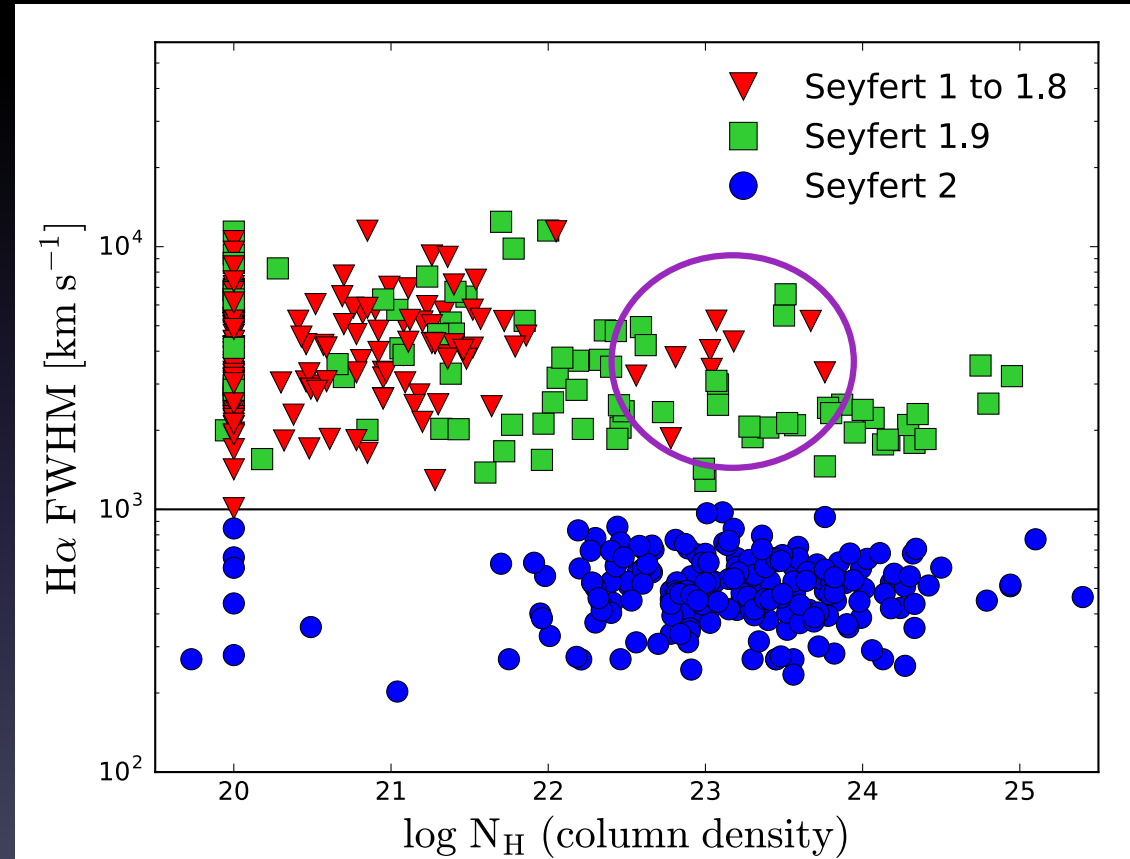
# **X-ray Obscured Type 1 AGN in BASS**

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**California Institute of Technology**  
**BASS Workshop 2019**



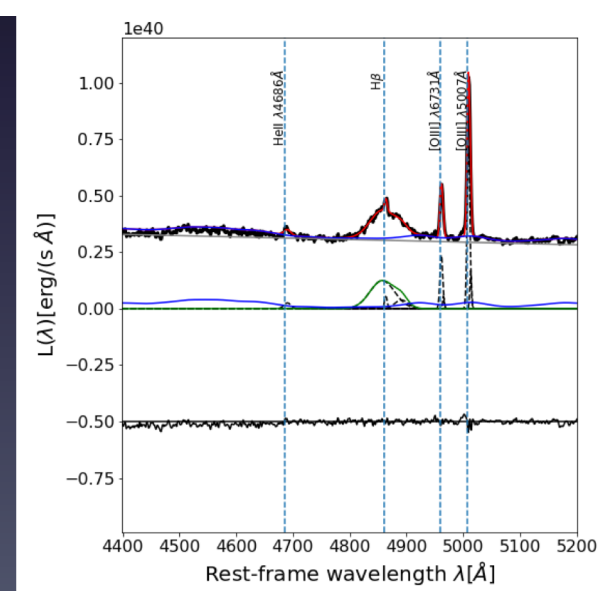
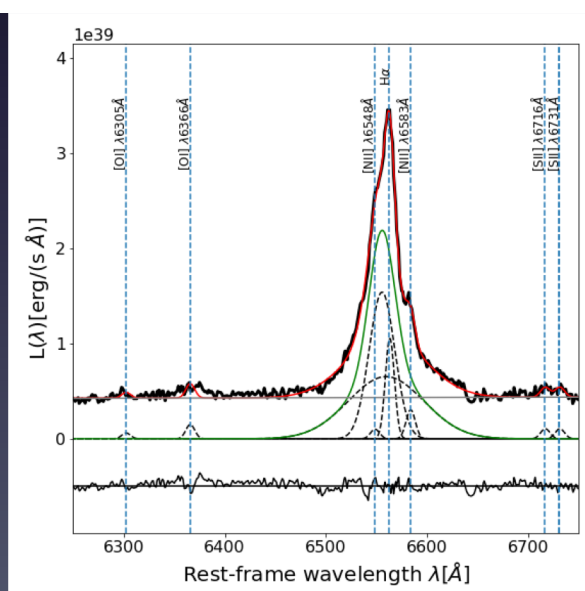
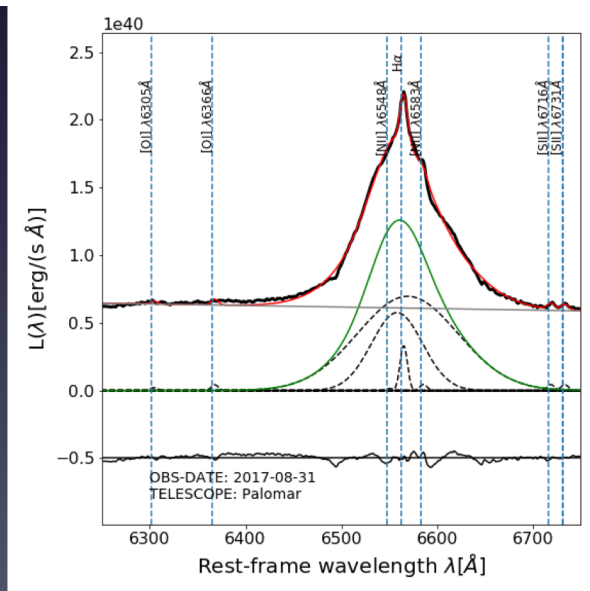
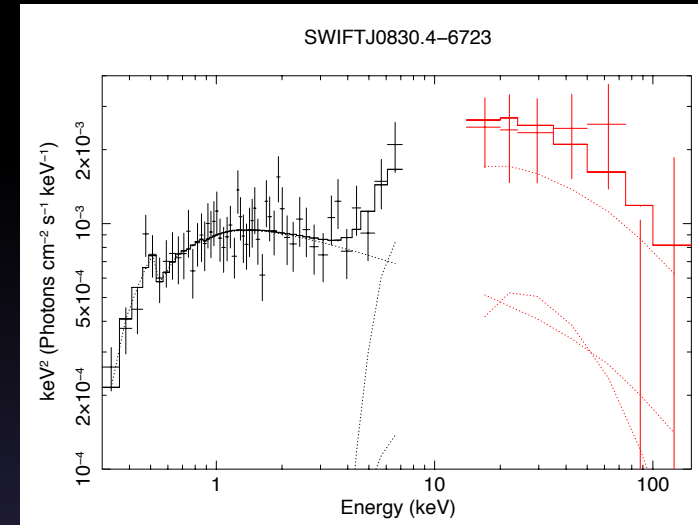
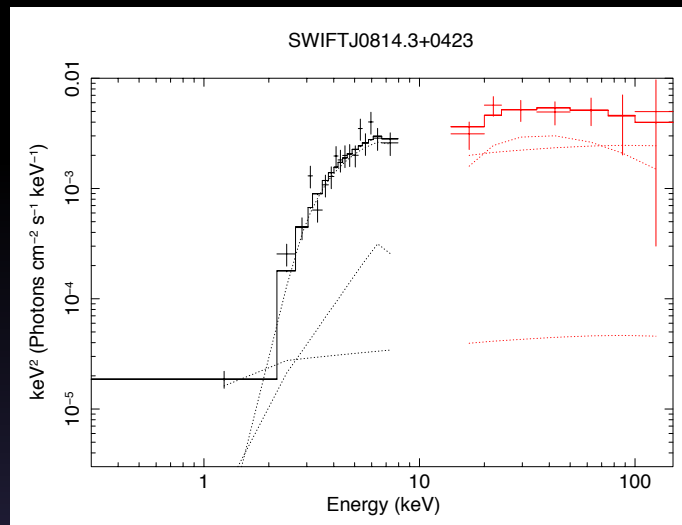
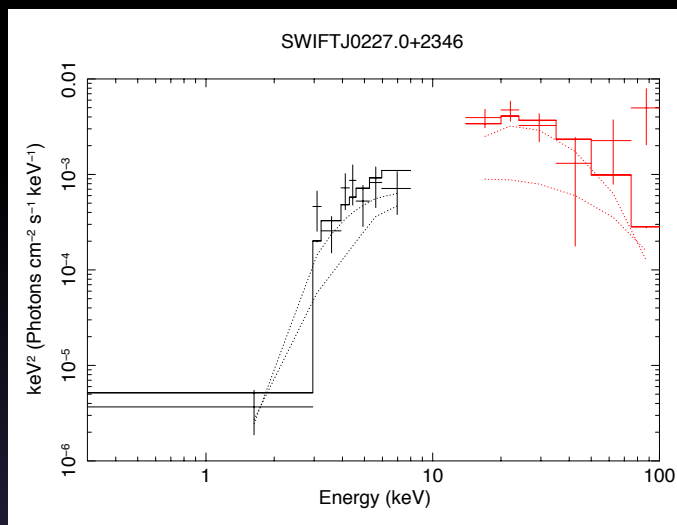
# OBSCURED TYPE 1 AGN IN BASS

- Optically classified as Type 1
- X-ray column density  $> 10^{22} \text{ cm}^{-2}$

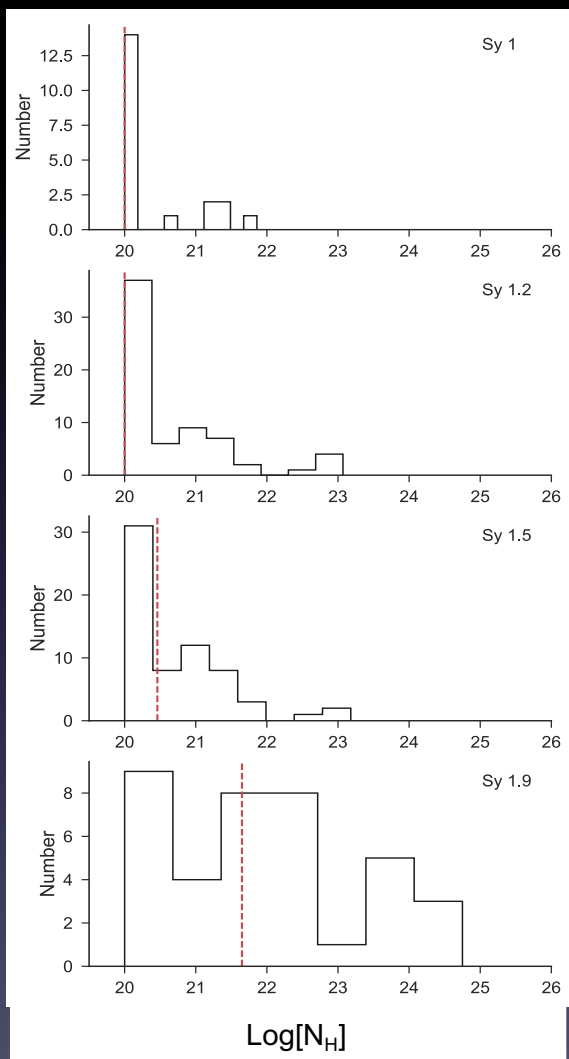


*Koss et al. 2017*

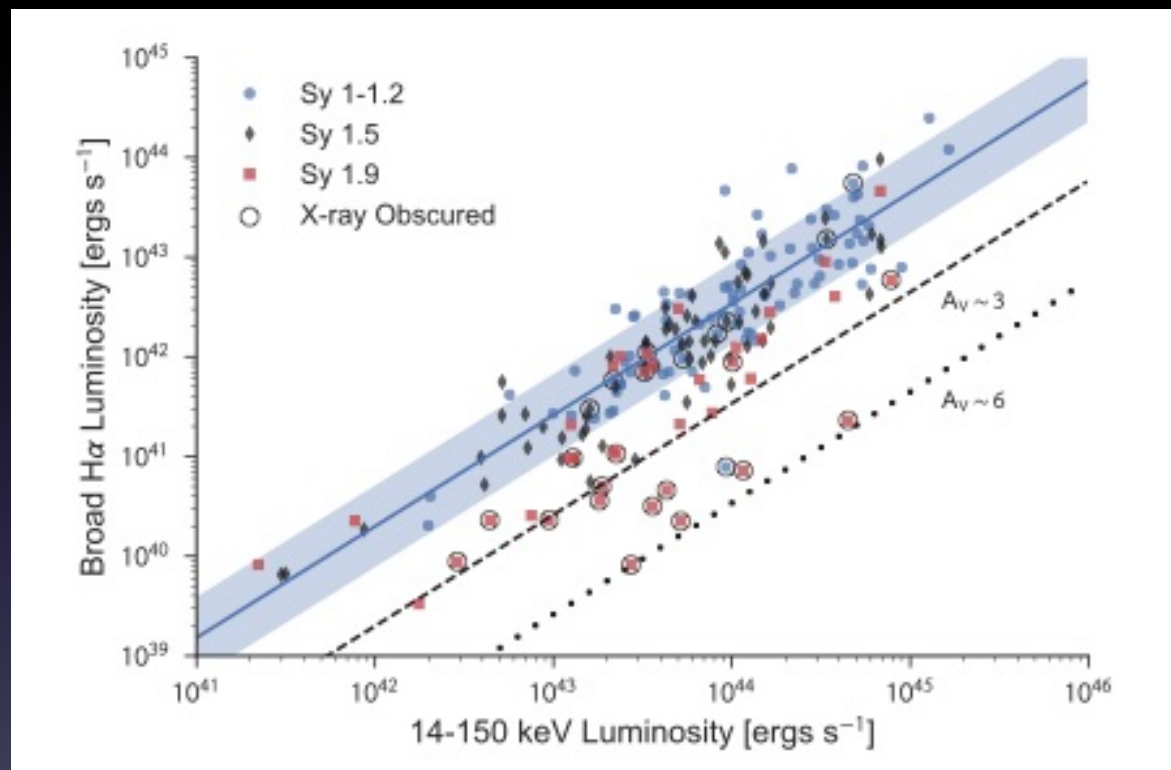
# OBSCURED TYPE 1 AGN IN BASS



# OBSCURED TYPE 1 AGN IN BASS

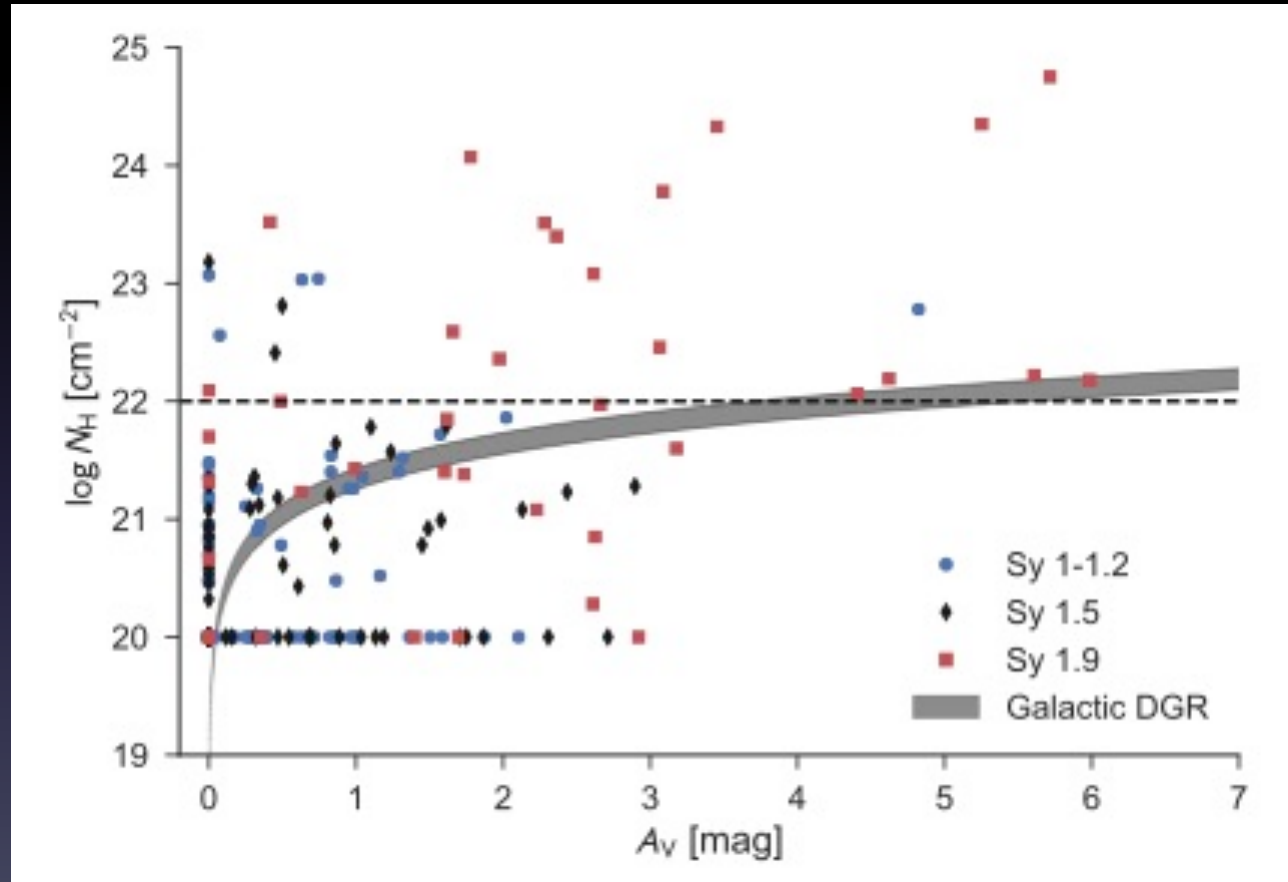


*Shimizu et al. 2017*



*Shimizu et al. 2017*

# OBSCURED TYPE 1 AGN IN BASS

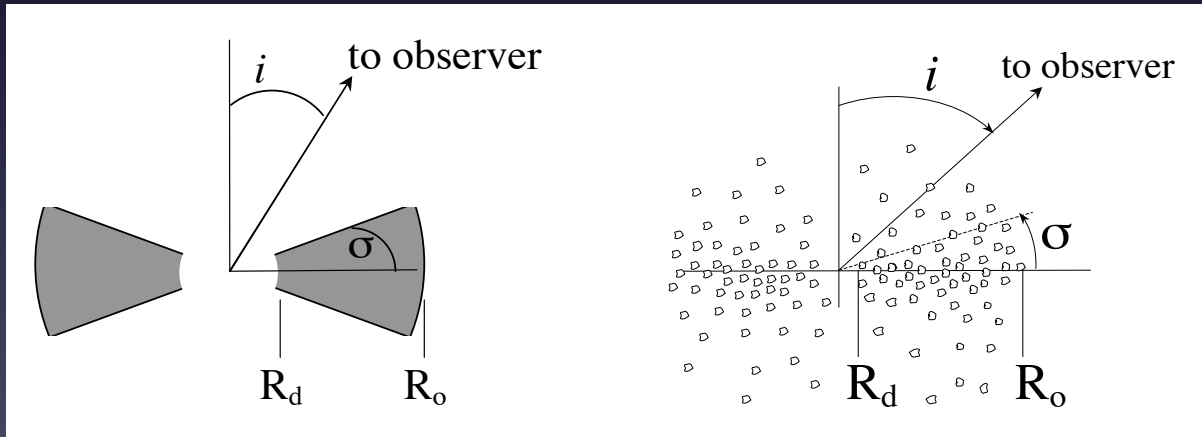


*Shimizu et al. 2017*

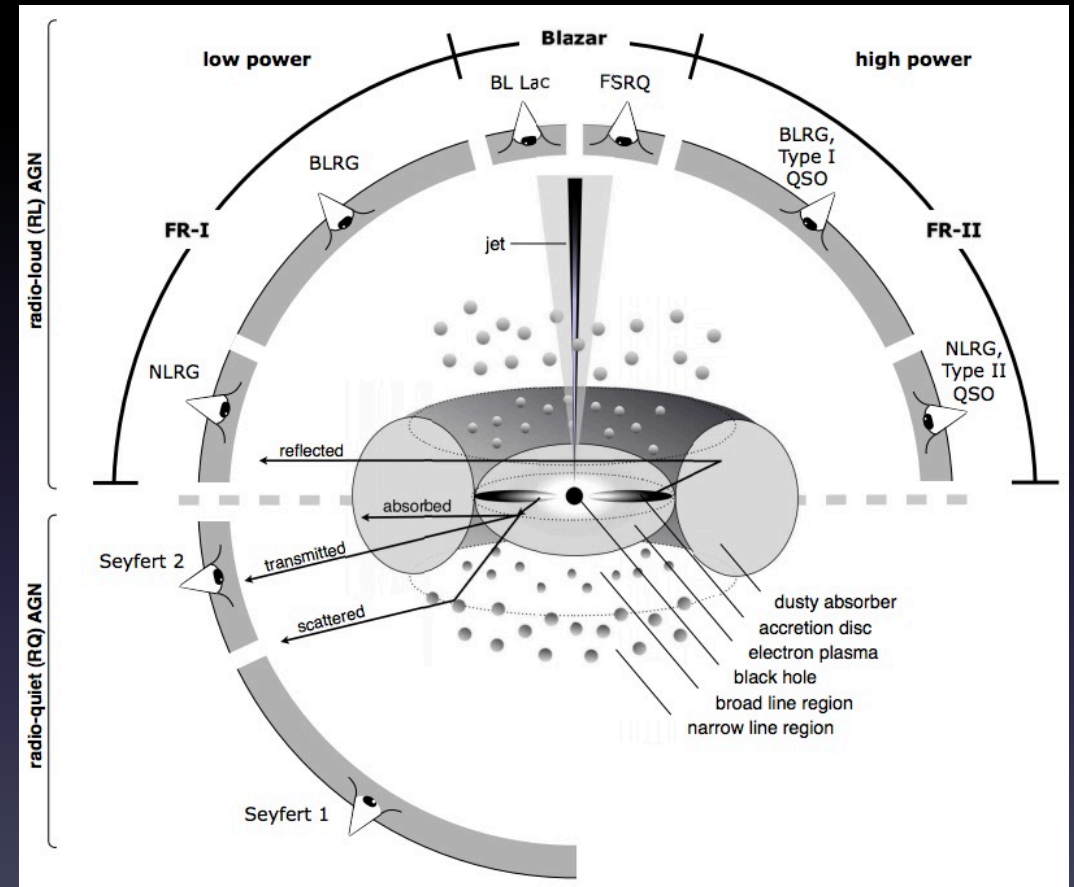
- ~14 % of Sy 1-1.9 from BASS sample are X-ray absorbed
- BLR provides extra obscuration towards corona

# PHYSICAL SCENARIOS

- Neutral, dust free torus
- Line of sight grazing edge of torus
- Cloud passing line of sight in clumpy torus
- Ionized gas outflows



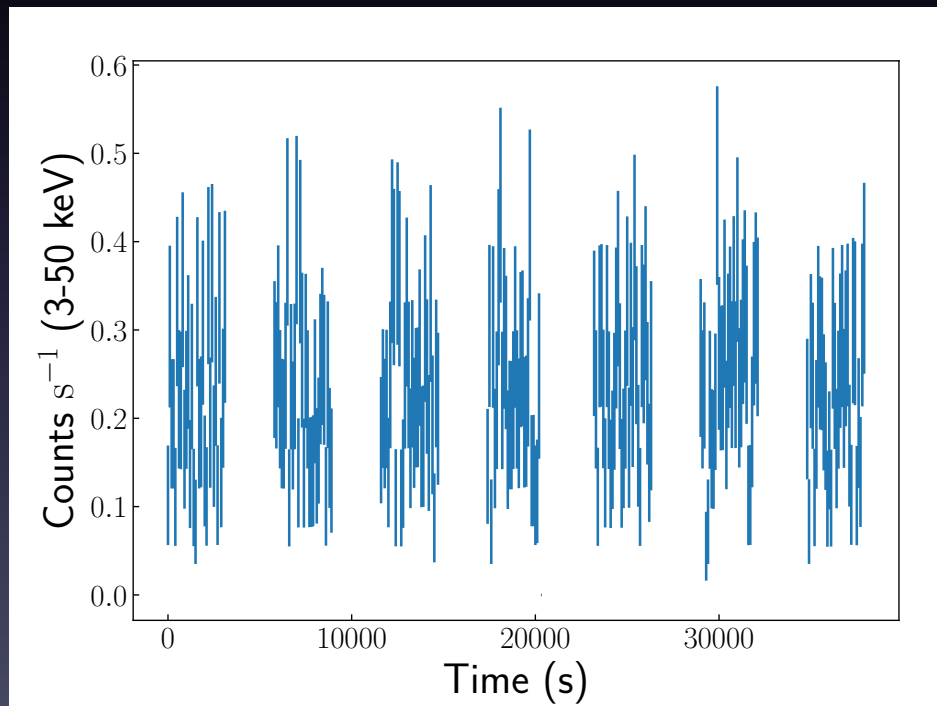
*M. Elitzur 2006*



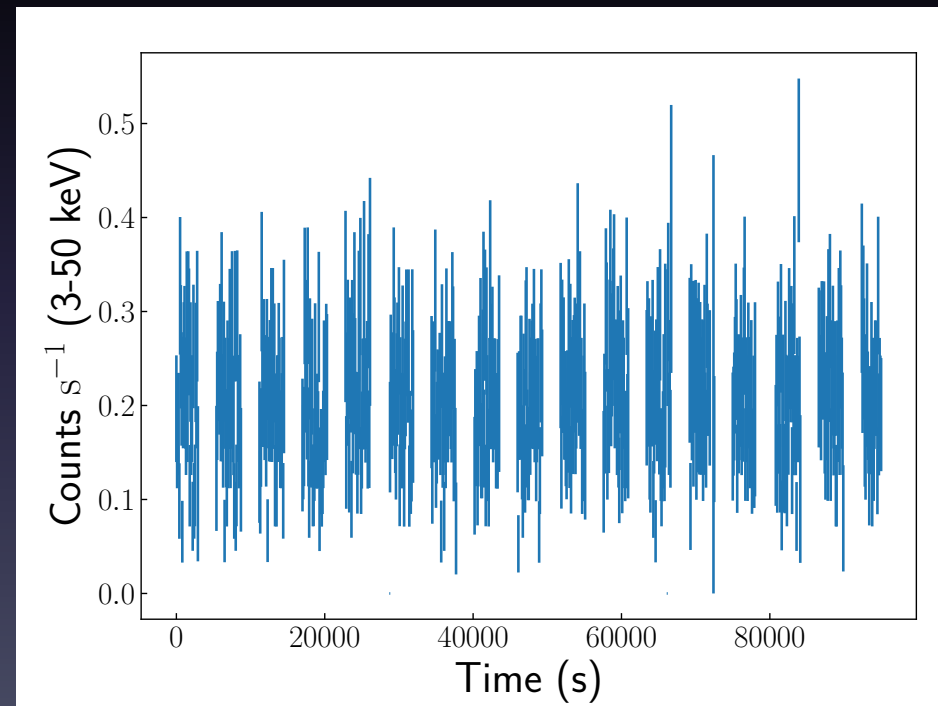
*Beckmann & Schrader 2012*

# NuSTAR OBSERVATIONS OF 2MASX J19301380+3410495

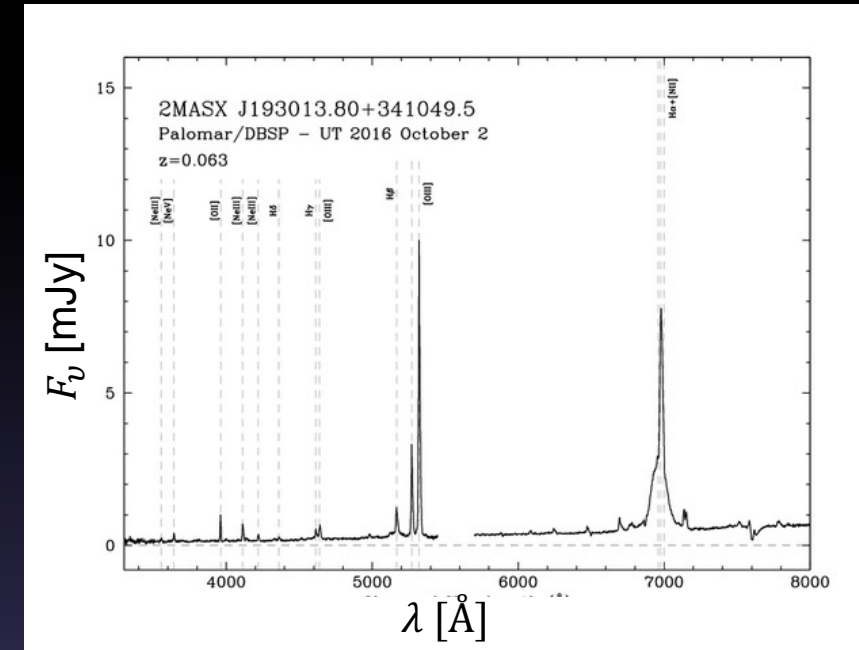
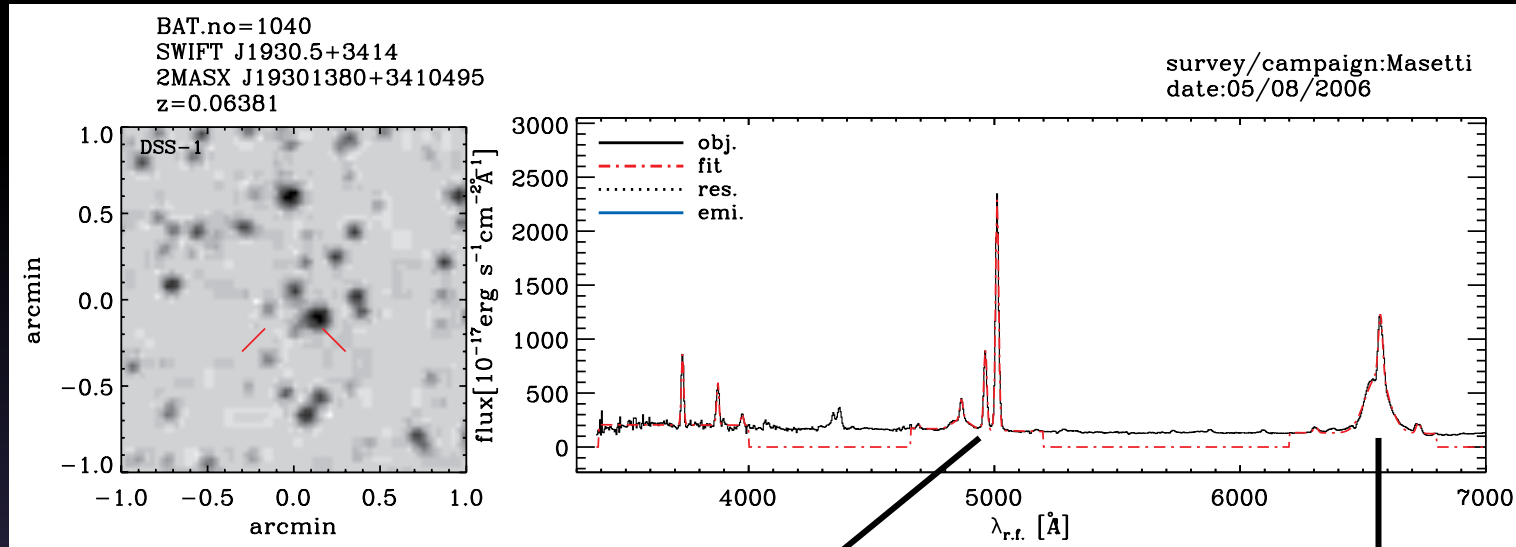
2016 obs, 20 ks



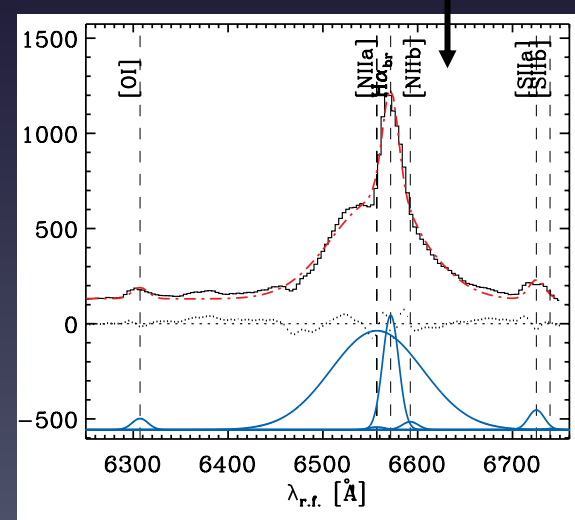
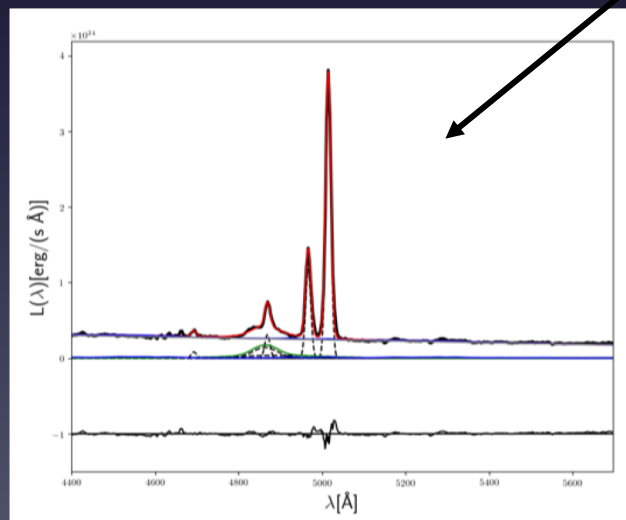
2017 obs, 50 ks



# OPTICAL SPECTRA

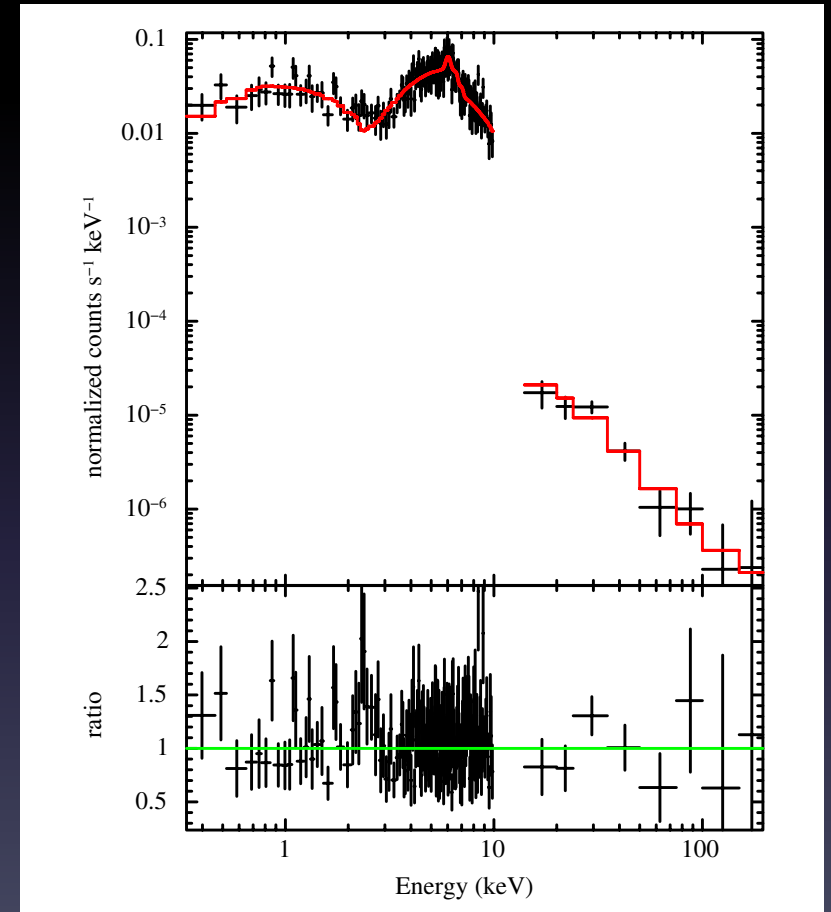


Optical spectra consistently  
show broad H $\alpha$  and H $\beta$  lines  $\rightarrow$   
Type 1 AGN



# X-RAY SPECTRAL MODELING

- Archival *Swift*/BAT, XRT, XMM spectra show  $N_H > 10^{23} \text{ cm}^{-2}$
- XSPEC models (this work):
  1. `zphabs x cabs x cutoffpl + pexrav + zgauss + (const. x cutoffpl)`
  2. `cabs x Tbfeo x cutoffpl + pexmon + (const. x cutoffpl)`
  3. `Borus`

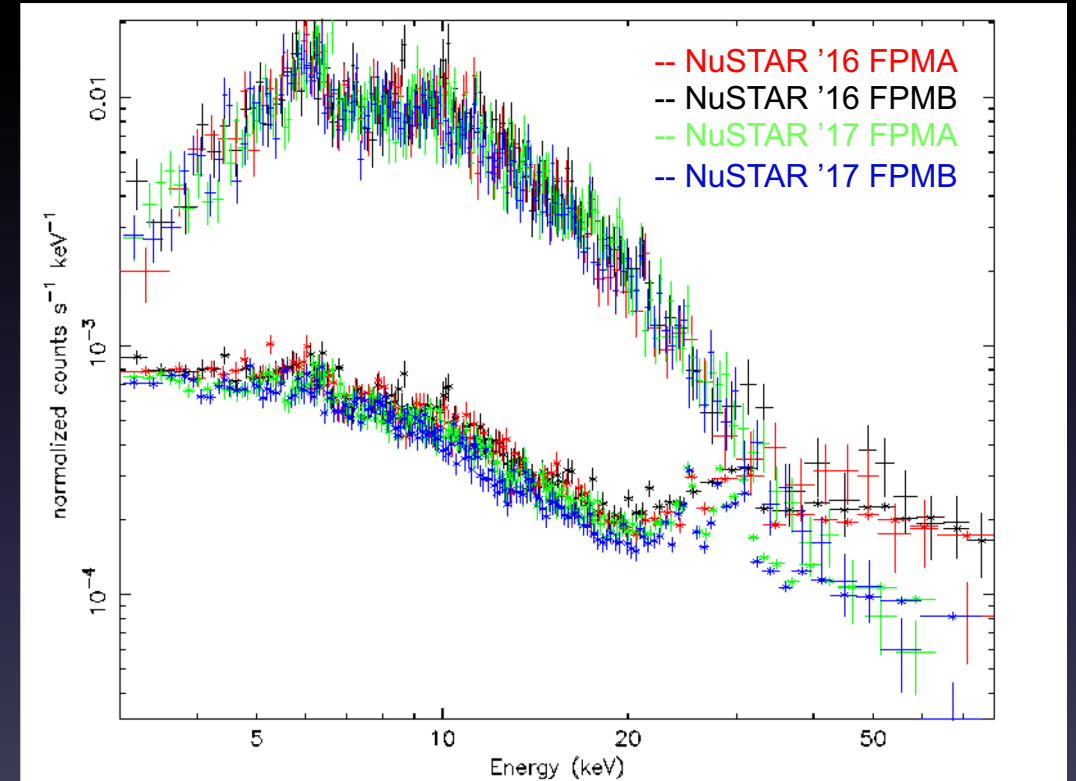


*Hogg et al. 2012*

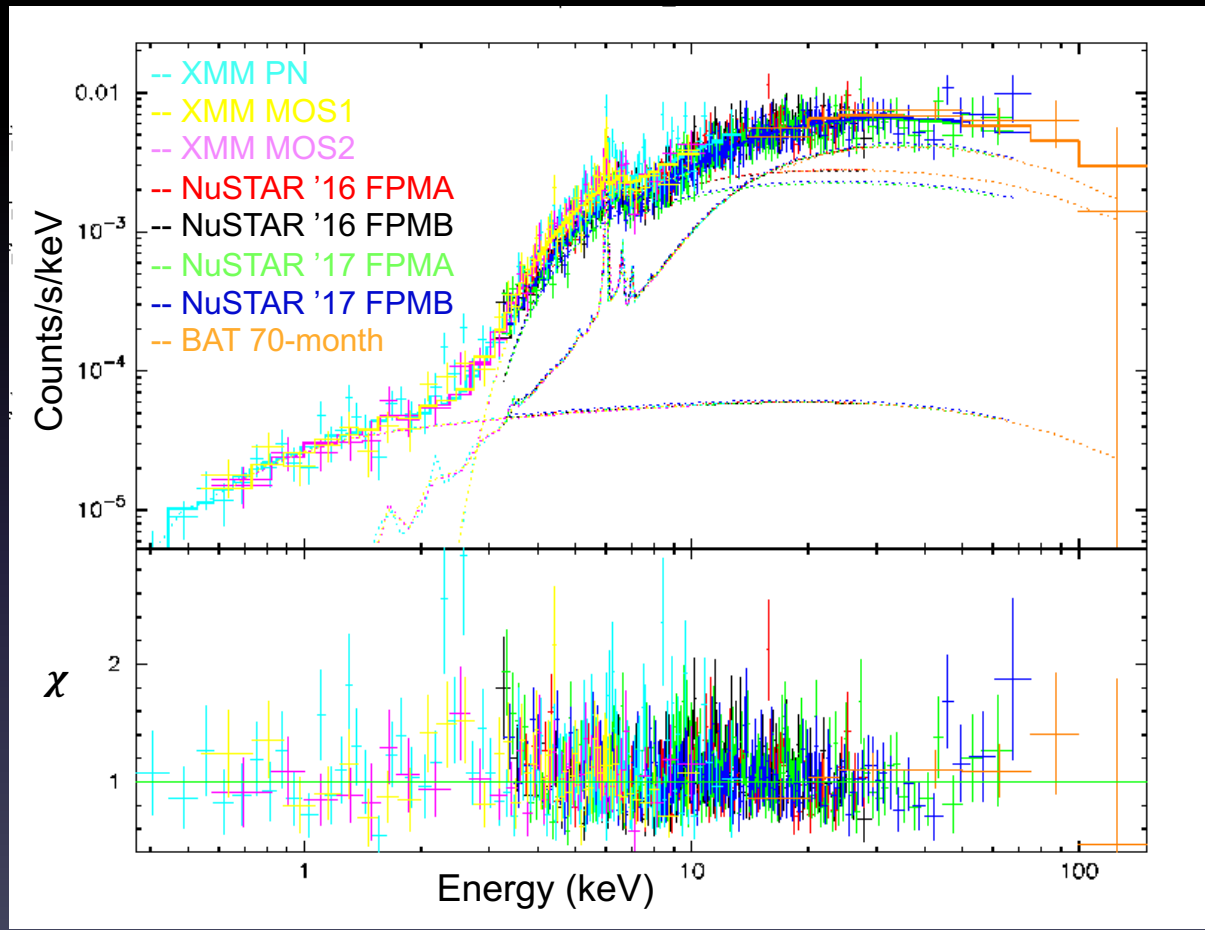
# X-RAY SPECTRAL MODELING

## Data:

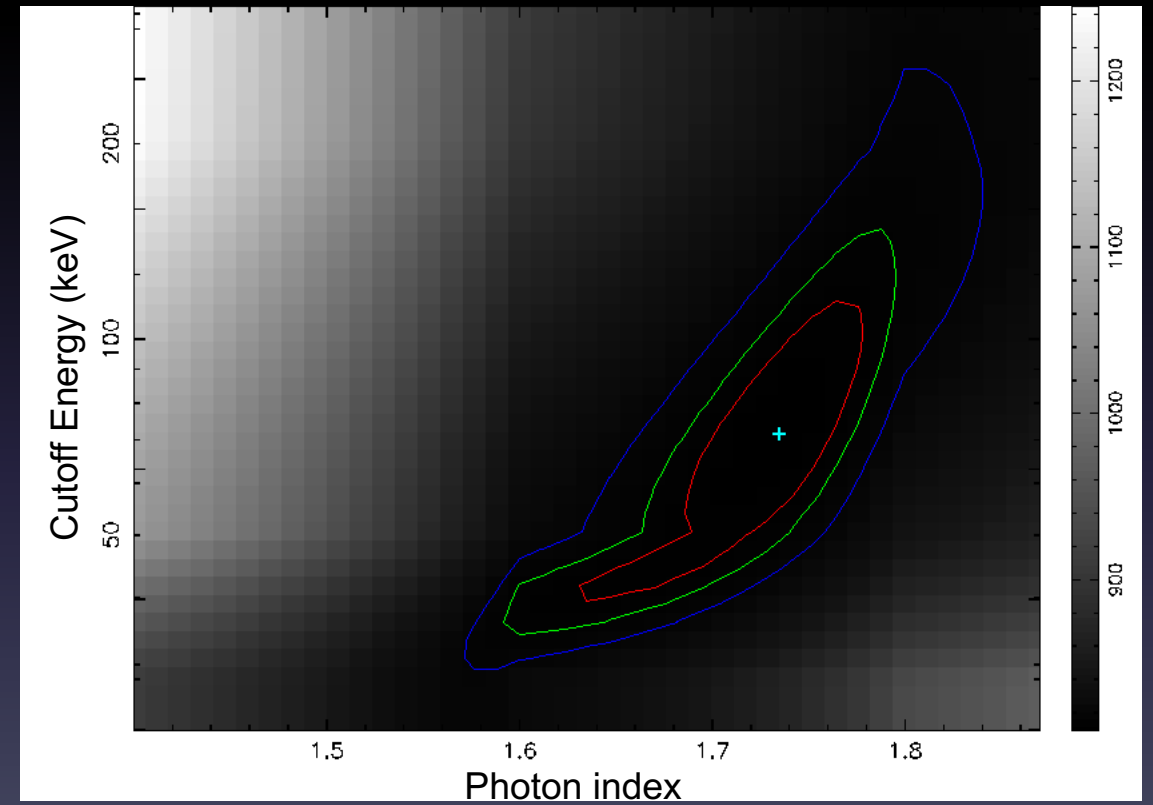
- NuSTAR – 2016 (up to 30 keV) + 2017 (3-79 keV)
- Archival XMM – 2009 (PN+MOS1+MOS2)
- Swift/BAT 70-month averaged spectra



# X-RAY SPECTRAL MODELING



Borus model



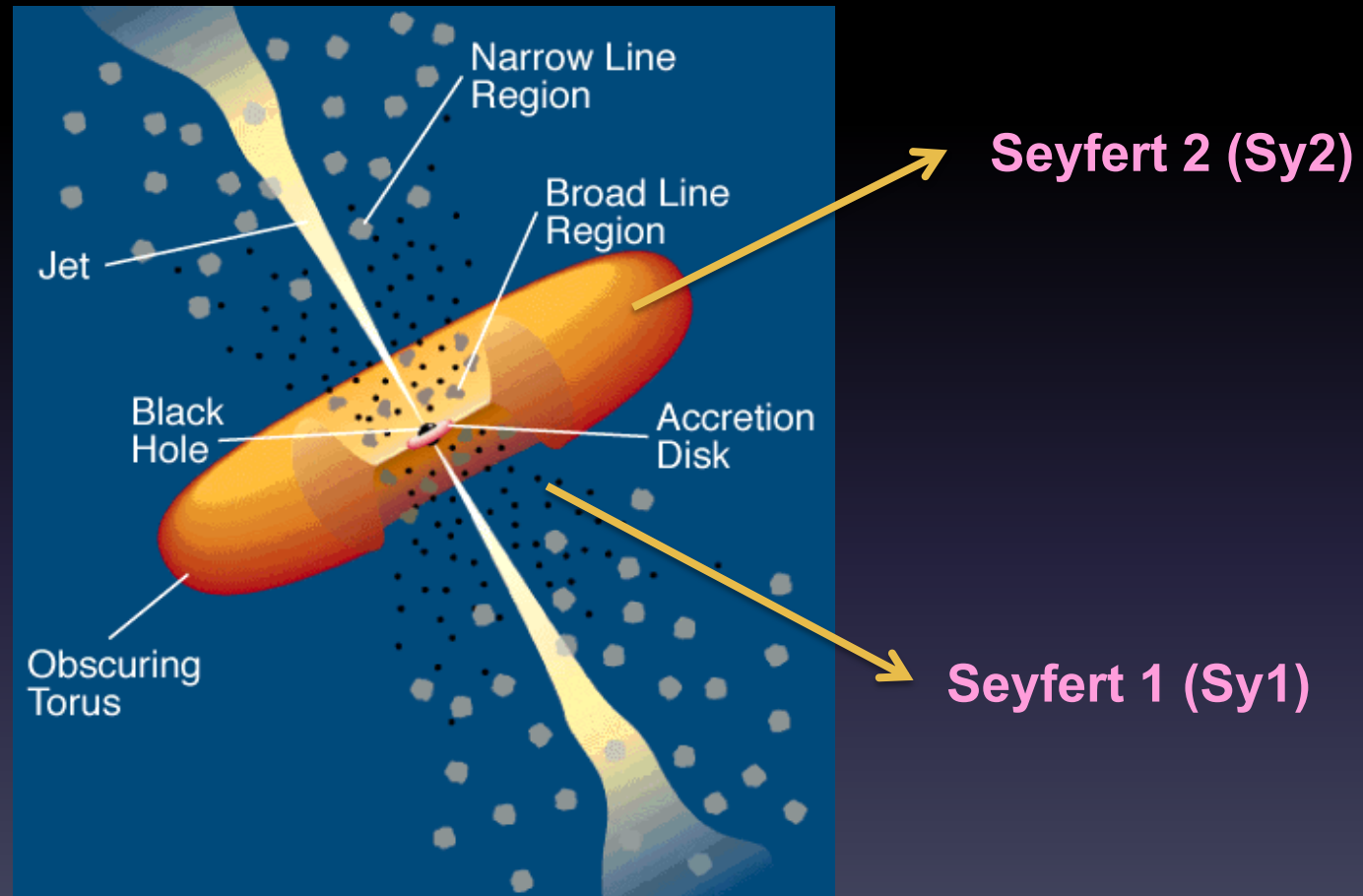
# X-RAY SPECTRAL MODELING

Model	$\chi^2/\text{dof}$	NuSTAR 2017 $N_{\text{H}}$ ( $\text{cm}^{-2}$ )	$\Gamma$	$E_{\text{cut}}$ (keV)	R	Fe abundance
pexrav	826/798	$(3.8 \pm 0.3) \times 10^{23}$	$1.35^{+0.18}_{-0.15}$	$49.9^{+19.0}_{-11.2}$	$-[0.90^{+0.23}_{-0.32}]$	1 (fixed)
pexmon	806/798	$5.5^{+0.7}_{-0.5} \times 10^{23}$	$1.33^{+0.21}_{-0.14}$	$49.9^{+38.3}_{-13.2}$	$-[0.70^{+0.10}_{-0.22}]$	1.09
borus	808/795	$4.9^{+0.5}_{-0.7} \times 10^{23}$	$1.73^{+0.05}_{-0.27}$	$71.9^{+24.2}_{-41.0}$	-	$0.45^{+0.06}_{-0.05}$

# FUTURE PLANS

- NuSTAR Cycle 5 proposal to study subsample of obscured type 1 AGN
- Currently available X-ray data: NuSTAR 20 ks legacy survey observations of some sources
- Ideas/thoughts/comments?

# AGN UNIFICATION



*Urry & Padovani 1995*